

Lecture 11: Beginning the Heat Equation

Model Problem: Heat flow in a Metal Rod

- As with modeling for the wave equation, we start with 1 dimension.
- Let $u(t, x)$ denote the temperature at time t of the rod at position x , for $x \in \mathbb{R}$ for now. We focus on two relationships: that between internal energy and external temp, and Fourier's law of heat conduction
- Thermal energy is proportional to a product of density & temperature

$$\text{Total Thermal Energy} = c \int_a^b \rho u dx \quad (\text{A})$$

↓ ↳ Specific heat ↳ Density (we assume constant)

- Fourier's law describes how heat moves:

$$q = -k \frac{\partial u}{\partial x} \quad (\text{B})$$

↳ Thermal flux ↳ Thermal conductivity constant

- Assume the rod is thermally isolated, so that conservation of energy dictates that energy change = flux difference

$$\frac{du}{dt}(t) = q(t, a) - q(t, b) = \int_a^b -\frac{\partial q}{\partial x} dx$$

While similarly, differentiating (A) gives

$$\frac{du}{dt} = c \int_a^b \rho \frac{\partial u}{\partial x} dx$$

Such that

$$\int_a^b c \rho \frac{\partial u}{\partial t} + \frac{\partial q}{\partial x} = 0$$

• as a & b were arbitrary, $C_P \frac{\partial U}{\partial T} + \frac{\partial Q}{\partial X} = 0$

By B)

$$\frac{\partial U}{\partial T} - \frac{K}{C_P} \frac{\partial^2 U}{\partial X^2} = 0 \quad (C)$$

the 1D Heat equation.

• If our rod is of finite length l , we impose B.C.

$$U(t, 0) = T_0 \quad U(t, l) = T_1 \quad (\text{Dirichlet B.C.})$$

to signify that we have fixed the temperature at the ends
 (holding in a ^{large} bath of water, for example)

These may be reduced to homogeneous ($= 0$) conditions
 by noting that

$$U_0(x) = T_0 \left(1 - \frac{x}{l}\right) + T_1 \frac{x}{l}$$

Satisfies the B.C. E(C). It is called the equilibrium
 Solution. By superposition, $U - U_0$ satisfies (C) and
 has $U(+, 0) = U(+, l) = 0$.

• Another possible case is having insulated ends, so no heat
 flows in or out.

$$\frac{\partial U}{\partial X}(t, 0) = \frac{\partial U}{\partial X}(t, l) = 0$$

(Neumann B.C.)

ex.) On the bounded interval $[0, \pi]$, we can find product solutions to the heat equation as in lemma

5.1 of lecture 9. For $u(0) = u(\pi) = 0$, Theorem 5.2 gives helmholtz solutions $\sin(nx)$, so our heat solutions are $u(t, x) = e^{-n^2 t} \sin(nx)$

Product equations:

$$\begin{aligned}\frac{dV}{dt} &= KV \\ \frac{d^2\Phi}{dx^2} &= K\Phi\end{aligned}$$

Notice $u(t, x) \rightarrow 0$ as $t \rightarrow \infty$ (losing heat ~~area~~ to approach 0 energy)

If we instead use insulated ends, we get solutions

$$u(t, x) = e^{-n^2 t} \cos(nx)$$

and $n=0$ yields a constant solution.

□

- The higher-dimensional heat eqn. may be derived

Similarly: $q = -K \nabla u$ and local conservation of energy is

$$CP \frac{\partial u}{\partial t} + \nabla \cdot q = 0$$

giving $\frac{\partial u}{\partial t} - \frac{K}{CP} \Delta u = 0$

⇒ This can be used to model Brownian Motion, which we will show now

•) Brownian Motion By Einstein's Argument:

- Suppose n particles are distributed on \mathbb{R} and in an interval of time τ , each particle's position changes by a random amount according to a distribution function Φ .
- The number of particles experiencing a displacement between G & $G+dG$ is

$$dn = n \Phi(G) dG$$

- Total # of particles is conserved: $\int_{\mathbb{R}} \Phi(G) dG = 1$

We also assume displacements to be symmetric in distribution:

$$\Phi(G) = \Phi(-G)$$

- Suppose the distribution of particles at time t is given by $\rho(t, x)$

By our displacement hypothesis,

$$\rho(t+\tau, x) = \int_{-\infty}^{\infty} \rho(t, x-G) \Phi(G) dG \quad (*)$$

To find an equation for ρ , Einstein takes the Taylor Expansion

$$\rho(t+\tau, x) = \rho(t, x) + \frac{\partial \rho}{\partial t}(t, x)\tau + \frac{1}{2} \frac{\partial^2 \rho}{\partial t^2}(t, x)\tau^2 \dots$$

and

$$\rho(t, x-G) = \rho(t, x) - \frac{\partial \rho}{\partial x}(t, x)G + \frac{1}{2} \frac{\partial^2 \rho}{\partial x^2}(t, x)G^2 \dots$$

Integrating the second,

$$\int_{-\infty}^{\infty} \rho(t, x-G) \Phi(G) dG = \rho(t, x) + \frac{1}{2} \frac{\partial^2 \rho}{\partial x^2}(t, x) \int_{-\infty}^{\infty} G^2 \Phi(G) dG \dots$$

Since $\int_{-\infty}^{\infty} \underbrace{G^{2K+1}}_{\text{odd}} \Phi(G) dG = 0$ for $K \in \mathbb{N}_0$

Then,

$$\rho(t, x) + \frac{\partial \rho}{\partial t}(t, x)\tau + \dots = \rho(t, x) + \frac{1}{2} \frac{\partial^2 \rho}{\partial x^2}(t, x) \int_{-\infty}^{\infty} G^2 \Phi(G) dG \dots$$

and keeping the leading term gives

$$\frac{\partial p}{\partial t}(t, x) \approx \frac{1}{2} \frac{\partial^2 p}{\partial x^2}(t, x) \int_{-\infty}^{\infty} \sigma^2 \phi(G) dG$$

- From Statistics, we also assume $D = \frac{1}{2} \int_{-\infty}^{\infty} \sigma^2 \phi(G) dG$ is constant, ~~known~~. So the equation for p becomes

$$\frac{\partial p}{\partial t} - D \frac{\partial^2 p}{\partial x^2} = 0.$$

\Rightarrow Diffusion model of particles.

Scale-Invariance Solution

- Consider the heat equation on \mathbb{R} , with physical constant 1

$$\frac{\partial u}{\partial t} - \frac{\partial^2 u}{\partial x^2} = 0 \quad (\text{H})$$

this will be a case of using "physical symmetries" to guess a solution. We notice that (H) is invariant under a rescaling $(t, x) \rightarrow (\lambda^2 t, \lambda x)$ for $\lambda \neq 0$, $\lambda \in \mathbb{R}$.

This suggests a change of variables to the scale-invariant

$y = \frac{x}{\sqrt{t}}$ might reduce (H) to an ODE.

- We try to find a solution $q(y) = u(t, x)$ for $t > 0$, so the chain rule gives

$$\frac{\partial u}{\partial t} = -\frac{y}{2t} q'$$

$$\frac{\partial^2 u}{\partial x^2} = \frac{y^2}{4t^2} q''$$

so (H) becomes $q'' = -\frac{y^2}{4t^2} q'$. Solving by sep. of. variables gives $q'(y) = q'(0) e^{-y^2/4}$ and $q(y) = q'(0) \int_0^y e^{-z^2/4} dz + q(0)$, or

$$u(t, x) = C_1 \int_0^{x/\sqrt{t}} e^{-y^2/4} dy + C_2$$

- Check that this satisfies (H)!

- To understand the situation $t \rightarrow 0$, note

$$\int_0^\infty e^{-s^2/4} ds = \sqrt{\pi}$$

(the Gaussian integral,
print available upon request)

thus

$$\lim_{t \rightarrow 0} u(t, x) = \begin{cases} C_1 \sqrt{\pi} + C_2 & x > 0 \\ 0 & x = 0 \\ -C_1 \sqrt{\pi} + C_2 & x < 0 \end{cases}$$

Therefore, we pick $C_1 = \frac{1}{\sqrt{4\pi}}$ $C_2 = \frac{1}{2}$ so

$$\tilde{U}(t, x) = \frac{1}{\sqrt{4\pi}} \int_0^{x/\sqrt{4t}} e^{-y^2/4} dy + \frac{1}{2}$$

has $\lim_{t \rightarrow 0} \tilde{U}(t, x) = \begin{cases} \frac{1}{2} & x > 0 \\ \frac{1}{2} & x = 0 \\ 0 & x < 0 \end{cases} = \Theta(x)$ "Heaviside Step Function"

- Why? Behind the scenes, we're using more complex theory to arrive at ~~more~~ convolutions. If we want initial condition $u(0, x) = \varphi(x) \in C_c^\infty(\mathbb{R})$, we observe

$$\int_{-\infty}^\infty \varphi'(z) \Theta(x-z) dz = \int_{-\infty}^x \varphi'(z) dz = \varphi(x)$$

so that we attempt to set

$$u(t, x) = \int_{-\infty}^\infty \varphi'(z) \tilde{U}(t, x-z) dz$$

Since \tilde{U} is C' in $t > 0$,

$$\begin{aligned} u(t, x) &= - \int_{-\infty}^\infty \varphi(z) \frac{\partial \tilde{U}}{\partial z}(t, x-z) dz \\ &\sim + \int_{-\infty}^\infty \frac{\varphi(z)}{\sqrt{4\pi t}} e^{--(x-z)^2/4t} dz \end{aligned}$$

Set $H_t(x) = \frac{1}{\sqrt{4\pi t}} e^{-x^2/4t}$ and

$$u(t, x) = \underbrace{\int_{-\infty}^\infty H_t(x-z) \varphi(z) dz}_{\text{Convolution}}$$

- We check this solution in the next lecture.